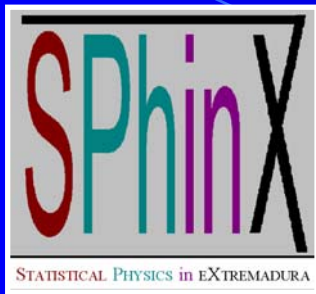


Granular Poiseuille flow



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Modelling and numerics of kinetic
dissipative systems (Lipari, May 31 -
June 4 2004)

Outline

- Gravity-driven Poiseuille flow for *conventional* gases.
- Newtonian description.
- Gravity-driven Poiseuille flow for *heated granular* gases.
- Kinetic theory description through second order in gravity.
- Results.
- Conclusions.

Jean-Louis Marie Poiseuille (1797-1869)



Poiseuille's law

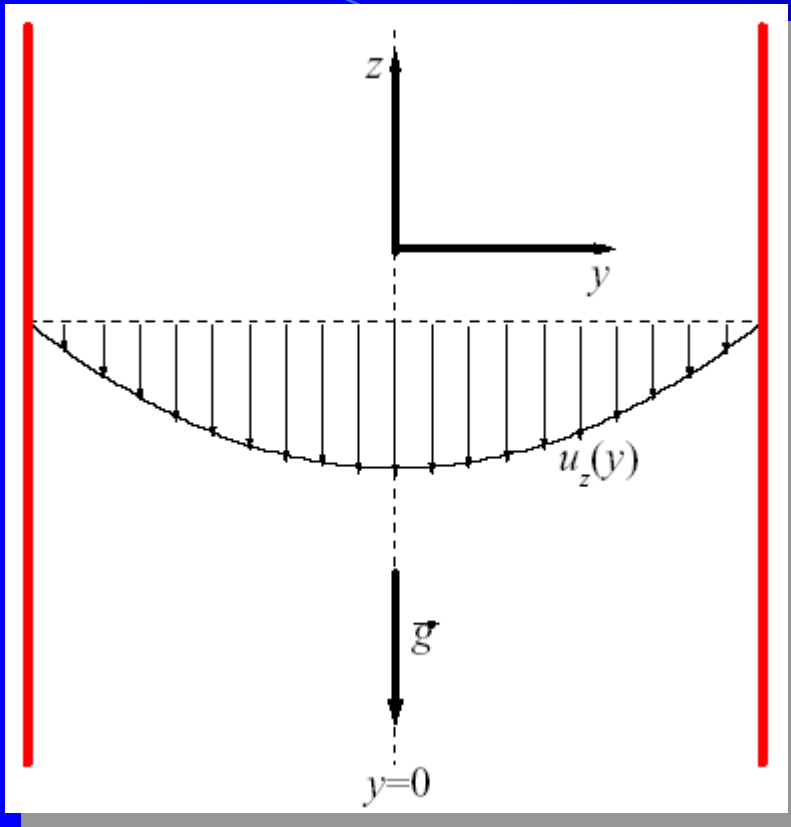
From Wikipedia, the free encyclopedia.

The **Poiseuille's law** (or the **Hagen-Poiseuille law** also named after [Gotthilf Heinrich Ludwig Hagen \(1797-1884\)](#) for his experiments in [1839](#)) is the [physical law](#) concerning the voluminal [laminar stationary](#) flow Φ_V of [incompressible](#) uniform [viscous](#) liquid (so called [Newtonian fluid](#)) through a cylindrical tube with the constant circular cross-section, experimentally derived in [1838](#), formulated and published in [1840](#) and [1846](#) by [Jean Louis Marie Poiseuille \(1797-1869\)](#), and defined by:

$$\Phi_V = \frac{dV}{dt} = v_s \pi r^2 = \frac{\pi r^4}{8\eta} \left(-\frac{dp^*}{dz} \right) = \frac{\pi r^4}{8\eta} \frac{\Delta p^*}{l},$$

where V is a volume of the liquid, poured in the time unit t , v_s median fluid [velocity](#) along the axial [cylindrical coordinate](#) z , r internal radius of the tube, Δp^* the pressure drop at the two ends, η dynamic fluid viscosity and l characteristic length along z , a linear dimension in a cross-section (in non-cylindrical tube).

Planar Poiseuille flow generated by a gravity field in a *conventional* gas



Conservation equations for momentum and energy

$$\frac{\partial P_{yy}}{\partial y} = 0$$

$$\frac{\partial P_{yz}}{\partial y} = -\rho g$$

$$P_{yz} \frac{\partial u_z}{\partial y} + \frac{\partial q_y}{\partial y} = 0$$

Navier-Stokes (Newtonian) description

$$P_{xx} = P_{yy} = P_{zz} = p$$

Equal normal stresses

$$P_{yz} = -\eta \frac{\partial u_z}{\partial y}$$

Newton's law

$$q_y = -\kappa \frac{\partial T}{\partial y}$$

Fourier's law

$$q_z = 0$$

No longitudinal heat flux

$$p(y) = p_0 = \text{const}$$

$$u_z(y) = u_0 + \frac{\rho_0 g}{2\eta_0} y^2 + \mathcal{O}(g^3)$$

$$T(y) = T_0 - \frac{\rho_0^2 g^2}{12\eta_0 \kappa_0} y^4 + \mathcal{O}(g^4)$$

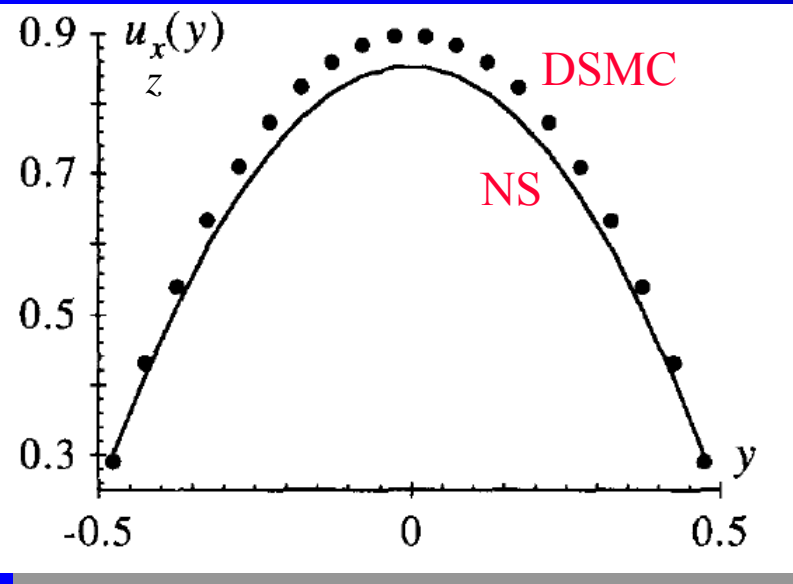
Temperature is *maximal* at the central layer ($y=0$)

Do NS predictions agree with computer simulations?

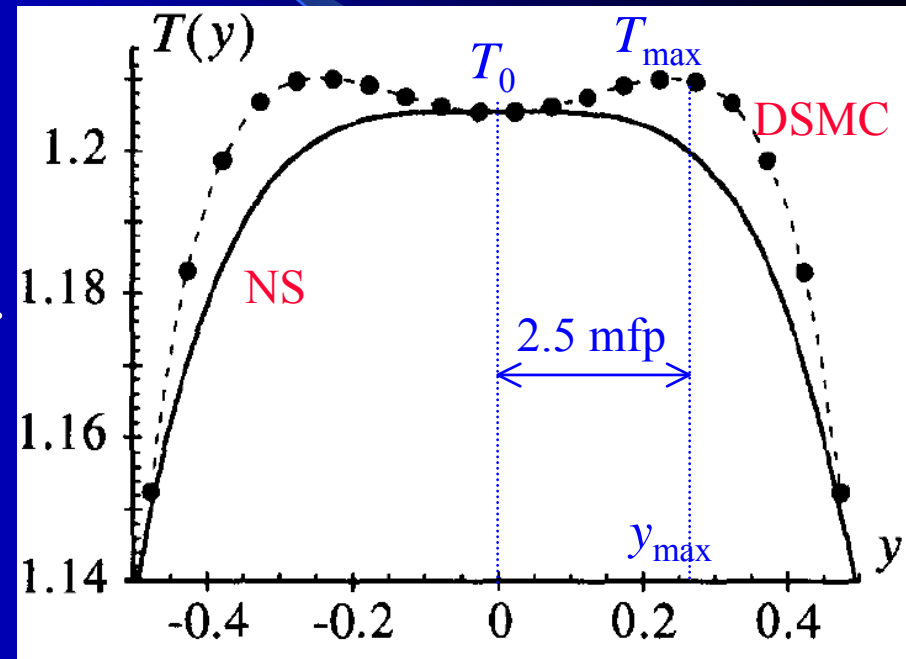
On the validity of hydrodynamics in plane Poiseuille flows

Physica A 240 (1997) 255–267

M. Malek Mansour^{a,*}, F. Baras^a, Alejandro L. Garcia^{b,1}



but ...



A Burnett-order effect?

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OCTOBER 1999

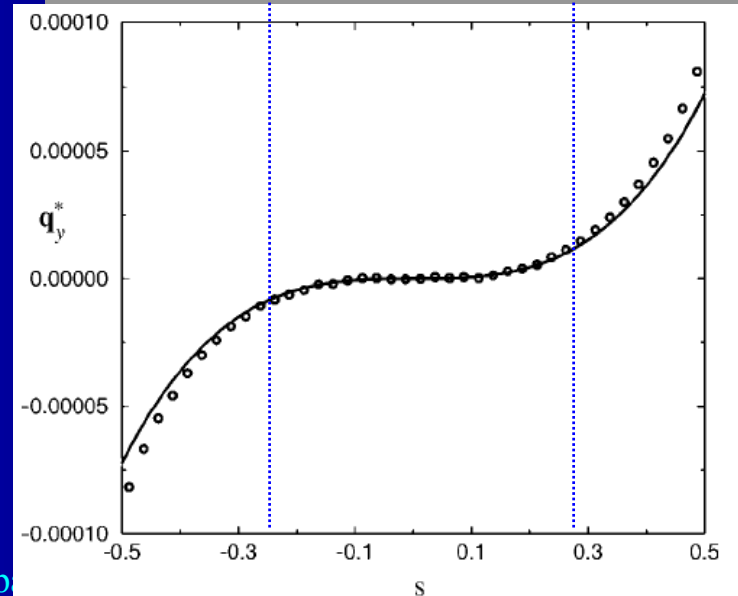
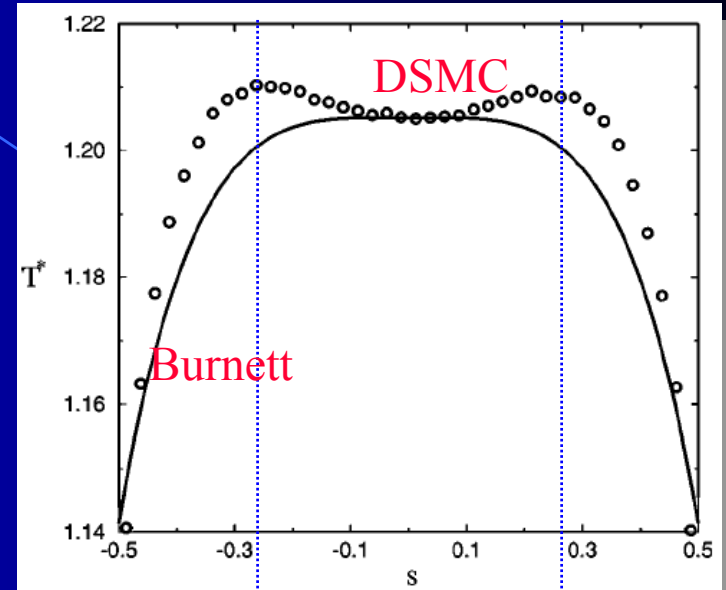
Burnett description for plane Poiseuille flow

F. J. Uribe Alejandro L. Garcia*

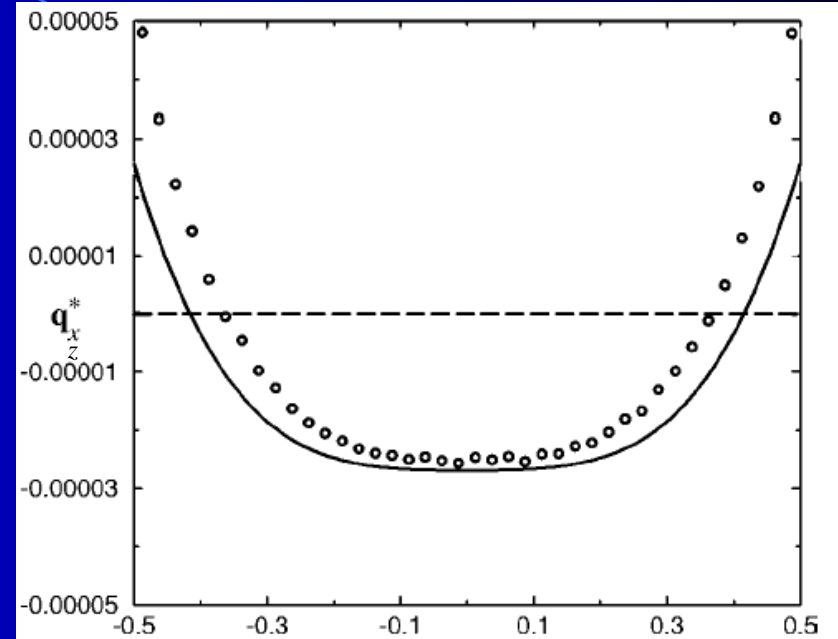
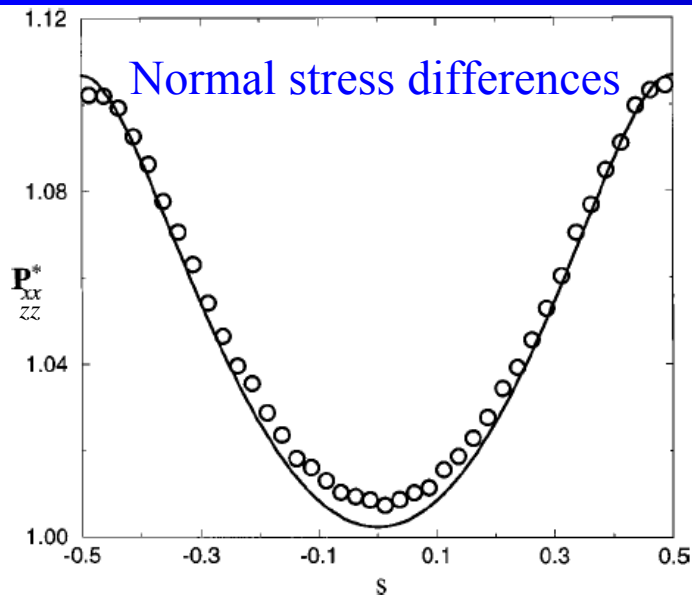
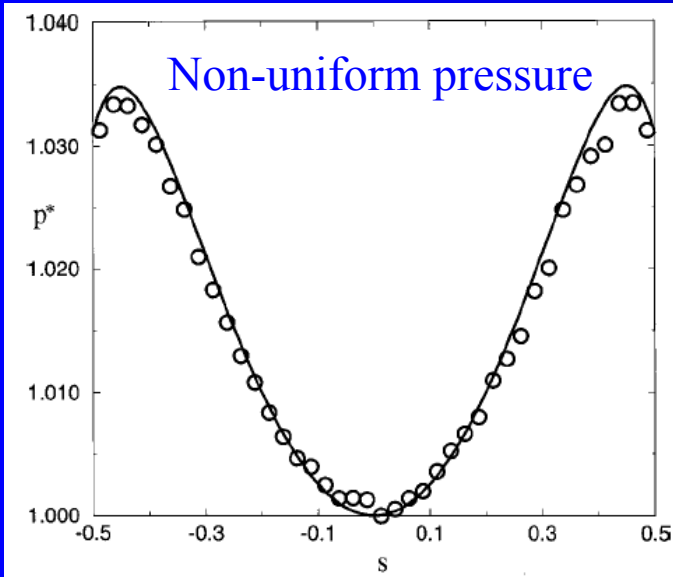
In the slab $y < |y_{\max}|$,

$$\text{sgn } q_y = \text{sgn } \partial T / \partial y$$

Heat flows from the colder
to the hotter layers!!



Other Non-Newtonian properties



Longitudinal component of the heat flux
(but no longitudinal thermal gradient!)

These Non-Newtonian effects are well accounted for by kinetic theory tools:

- Perturbative solution of the BGK and Boltzmann-Maxwell kinetic equations (M. Tij, M. Sabbane, A.S.).
- Grad's method applied to the Boltzmann equation for hard spheres (S. Hess, M. Malek Mansour, D. Risso, P. Cordero).
- Asymptotic analysis of the BGK model for small Knudsen numbers (K. Aoki, S. Takata, T. Nakanishi).

Is the gravity-driven Poiseuille flow relevant to real gases?

$$\frac{T_{\max} - T_0}{T_0} \gtrsim 10^{-2} \Rightarrow g \frac{\lambda}{v_{\text{th}}^2} \gtrsim 2 \times 10^{-2}$$

λ : mean free path; v_{th} : thermal velocity

Argon at room conditions $\left\{ \begin{array}{l} g=9.8 \text{ m/s}^2 \\ \lambda \sim 700 \text{ \AA} \\ v_{\text{th}} \sim 400 \text{ m/s} \end{array} \right\} \boxed{g\lambda/v_{\text{th}}^2 \sim 10^{-12} !!}$

Fluidized *granular* particles



They are *mesoscopic* particles ($\sigma \sim 1$ mm)

Some typical values

$$\left\{ \begin{array}{l} g = 9.8 \text{ m/s}^2 \\ \lambda \approx 1 \text{ mm} - 1 \text{ cm} \\ v_{\text{th}} \gtrsim 1 \text{ m/s} \end{array} \right\} \quad \boxed{g\lambda/v_{\text{th}}^2 \sim 10^{-3} - 10^{-1}}$$

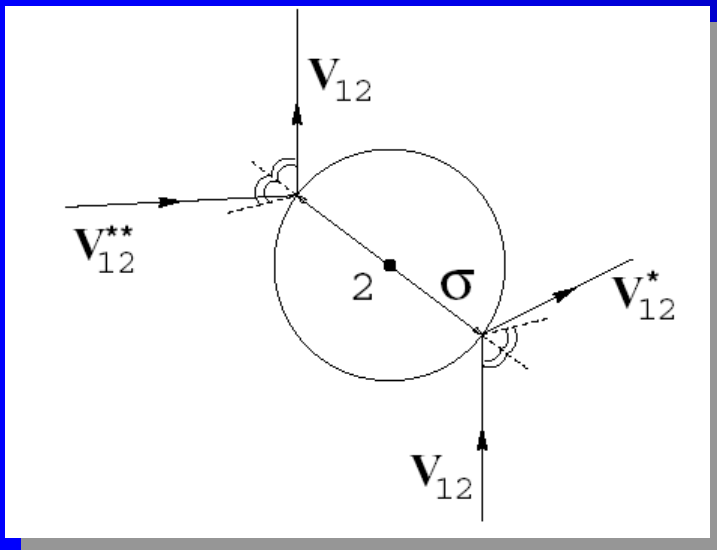
The dimensionless parameter $g\lambda/v_{\text{th}}^2$ measures the strength of gravity between collisions. It can be:

- Large enough as to produce measurable effects.
- Small enough as to allow for a perturbative treatment.

Our main goal is:

- Call attention to the fact that non-Newtonian properties in the gravity-driven Poiseuille flow can be observable on granular gases under laboratory conditions.
- Assess the influence of inelasticity on the hydrodynamic fields and their fluxes.
E.g., is $(T_{\max} - T_0)/T_0$ enhanced or inhibited by inelasticity?

A gas of (smooth) *inelastic* hard spheres



(After T.P.C. van Noije & M.H. Ernst)

α : coefficient of (normal) restitution

Direct collision

$$\mathbf{v}'_1 = \mathbf{v}_1 - \frac{1+\alpha}{2}(\mathbf{v}_{12} \cdot \hat{\boldsymbol{\sigma}})\hat{\boldsymbol{\sigma}}, \quad \mathbf{v}'_2 = \mathbf{v}_2 + \frac{1+\alpha}{2}(\mathbf{v}_{12} \cdot \hat{\boldsymbol{\sigma}})\hat{\boldsymbol{\sigma}}$$

Restituting collision

$$\mathbf{v}''_1 = \mathbf{v}_1 - \frac{1+\alpha}{2\alpha}(\mathbf{v}_{12} \cdot \hat{\boldsymbol{\sigma}})\hat{\boldsymbol{\sigma}}, \quad \mathbf{v}''_2 = \mathbf{v}_2 + \frac{1+\alpha}{2\alpha}(\mathbf{v}_{12} \cdot \hat{\boldsymbol{\sigma}})\hat{\boldsymbol{\sigma}}$$

Boltzmann equation

$$\left(\partial_t + \mathbf{v} \cdot \nabla + \underbrace{\mathbf{g} \cdot \frac{\partial}{\partial \mathbf{v}}}_{\text{Gravity}} + \underbrace{\mathcal{F}}_{\text{External driving}} \right) f = \underbrace{J[f, f]}_{\text{Inelastic collisions}}$$

Gravity

External
driving

Inelastic
collisions

$$J[f, f] = \sigma^2 \int d\mathbf{v}_1 \int d\hat{\boldsymbol{\sigma}} \Theta((\mathbf{v} - \mathbf{v}_1) \cdot \hat{\boldsymbol{\sigma}}) [(\mathbf{v} - \mathbf{v}_1) \cdot \hat{\boldsymbol{\sigma}}] [\alpha^{-2} f(\mathbf{v}'') f(\mathbf{v}_1'') - f(\mathbf{v}) f(\mathbf{v}_1)]$$

Collisional “cooling”

$$\frac{m}{3} \int d\mathbf{v} V^2 J[f, f] = -\zeta n T$$

Cooling rate

External “heating”
(e.g., vibrations)

$$\frac{m}{3} \int d\mathbf{v} V^2 \mathcal{F} f(\mathbf{v}) = -\gamma n T$$

Heating rate

$\mathbf{V} \equiv \mathbf{v} - \mathbf{u}$ (peculiar velocity)

$$\zeta \simeq \nu \frac{5}{12} (1 - \alpha^2)$$

Gaussian approximation

$$\nu = \frac{16}{5} n \sigma^2 \left(\frac{\pi T}{m} \right)^{1/2}$$

Effective collision frequency

White noise driving

It is a *bulk* heating mechanism that intends to mimic the effect of boundary driving (e.g., vibrations).

Each particle is subjected to the action of a stochastic force with white noise properties:

$$\langle \mathbf{F}^{\text{wn}}(t) \rangle = \mathbf{0}, \quad \langle F_{\alpha}^{\text{wn}}(t) F_{\beta}^{\text{wn}}(t') \rangle = m^2 \xi^2 \delta_{\alpha\beta} \delta(t - t')$$

During a small time step Δt , each particle receives a “kick,” so its velocity is incremented by a random amount $\Delta \mathbf{v}$

$$\Delta t \Rightarrow |\Delta \mathbf{v}| \sim \xi \sqrt{\Delta t}$$

Diffusion in velocity space: $\mathcal{F} = -\frac{\xi^2}{2} \left(\frac{\partial}{\partial \mathbf{v}} \right)^2 \Rightarrow \gamma = \frac{m\xi^2}{T}$

Heating rate

Our choice: The white noise compensates *locally* for the collisional cooling.

$$\gamma = \zeta \Rightarrow \frac{|\Delta \mathbf{v}|}{v_{\text{th}}} \sim \sqrt{\nu \Delta t (1 - \alpha^2)}$$

The *relative* magnitude of the kick scales with (the square root of) the (local) probability of a collision.

Associated NS transport coefficients:

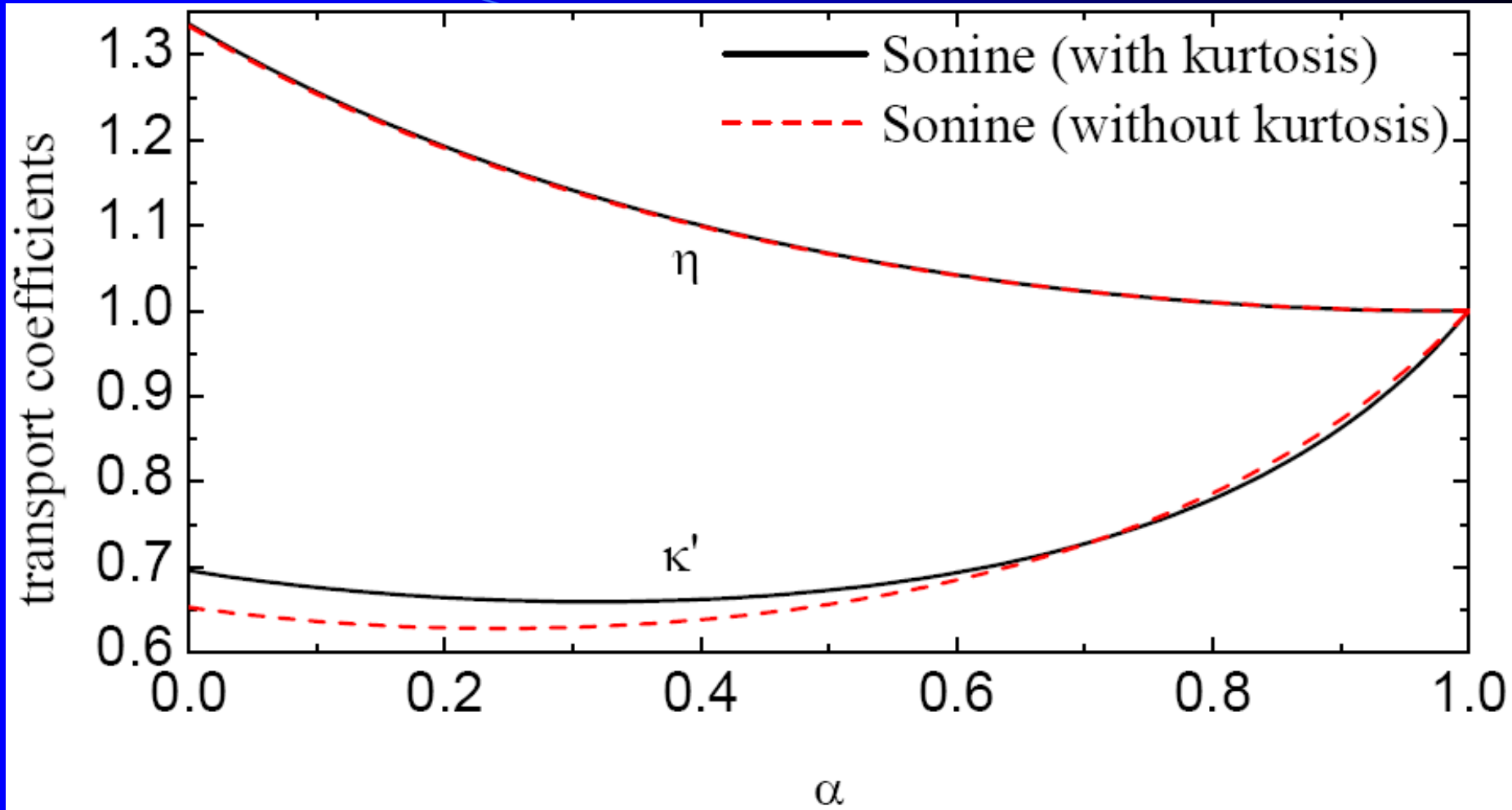
(Garzó & Montanero, 2002)

$$\eta \simeq \frac{p}{\nu} \frac{4}{(1 + \alpha)(3 - \alpha)}, \quad \kappa \simeq \frac{5p}{2m\nu} \frac{48}{(1 + \alpha)(49 - 33\alpha)}$$

Increases with inelasticity

Decreases with inelasticity ($\alpha \gtrsim 0.4$)

Increases with inelasticity ($\alpha \lesssim 0.4$)



Stationary Boltzmann equation

$$\left(\underbrace{-\frac{\zeta T}{2m} \frac{\partial^2}{\partial \mathbf{v}^2}}_{\text{White noise heating}} - \underbrace{g \frac{\partial}{\partial v_z}}_{\text{Gravity}} + v_y \frac{\partial}{\partial y} \right) f = \underbrace{J[f, f]}_{\text{Inelastic collisions}}$$

White noise heating

Gravity

Inelastic collisions

BGK-like kinetic model:
(Brey, Dufty, A.S.)

$$J[f, f] \rightarrow -\underbrace{\beta(\alpha)\nu}_{\text{Modified collision frequency}}(f - f_\ell) + \underbrace{\frac{\zeta}{2} \frac{\partial}{\partial \mathbf{v}} \cdot [(\mathbf{v} - \mathbf{u}) f]}_{\text{Effective drag force: mimics cooling}}$$

Modified collision frequency

Effective drag force:
mimics cooling

$$f_\ell(\mathbf{r}, \mathbf{v}; t) = n(\mathbf{r}, t) \left[\frac{m}{2\pi T(\mathbf{r}, t)} \right]^{3/2} \exp \left[-\frac{m(\mathbf{v} - \mathbf{u}(\mathbf{r}, t))^2}{2T(\mathbf{r}, t)} \right]$$

Local Gaussian distribution

Digression: How reliable is the BGK-like model?

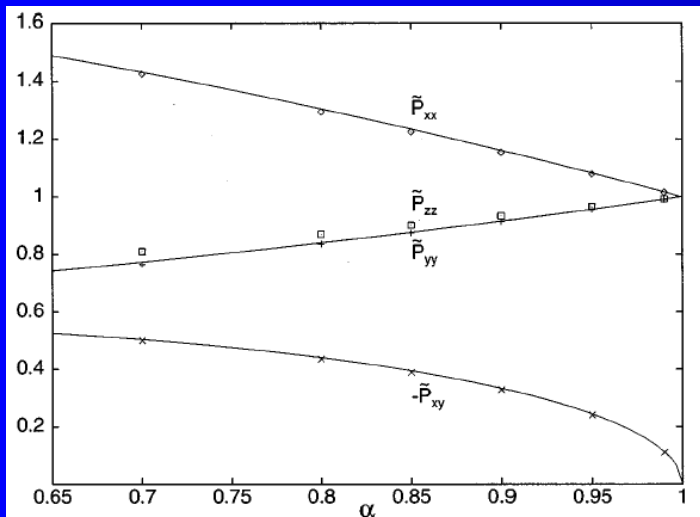
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MARCH 1997

Steady uniform shear flow in a low density granular gas

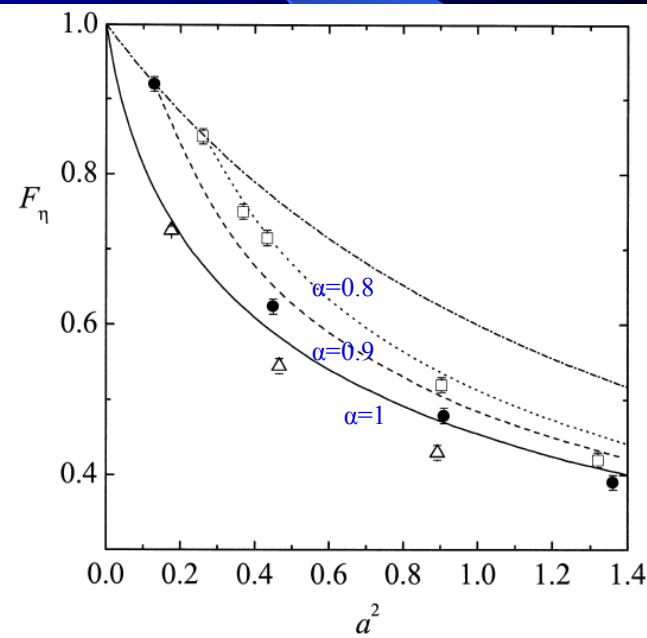
J. J. Brey, M. J. Ruiz-Montero, and F. Moreno



Journal of Statistical Physics, Vol. 103, Nos. 5/6, 2001

Nonlinear Couette Flow in a Low Density Granular Gas

M. Tij,¹ E. E. Tahiri,² J. M. Montanero,³ V. Garzó,⁴ A. Santos,⁵
and J. W. Dufty⁵



Perturbation expansion

$$f(y, \mathbf{V}) = f_\ell(y, \mathbf{V}) \left[1 + \Phi^{(1)}(y, \mathbf{V})g + \Phi^{(2)}(y, \mathbf{V})g^2 + \mathcal{O}(g^3) \right]$$

Velocity distribution function

$$p(y) = p_0 + p^{(2)}(y)g^2 + \mathcal{O}(g^4)$$

$$u_z(y) = u_0 + u^{(1)}(y)g + \mathcal{O}(g^3)$$

$$T(y) = T_0 + T^{(2)}(y)g^2 + \mathcal{O}(g^4)$$

Hydrodynamic profiles

Structure of the solution through second order:

$$\Phi^{(1)}(y, \mathbf{V}) = V_z(a_0 + a_1V_y^2 + a_2V_yy)$$

$$\begin{aligned} \Phi^{(2)}(y, \mathbf{V}) = & b_0 + b_1V_y^2 + b_2V_yy + b_3y^2 + b_4V_y^4 + b_5V_y^3y + b_6V_y^2y^2 + b_7V_yy^3 \\ & + (c_0 + c_1V_y^2 + c_2V_yy + c_3y^2 + c_4V_y^4 + c_5V_y^3y + c_6V_y^2y^2) V_z^2 \\ & + (d_0 + d_1V_y^2 + d_2V_yy + d_3y^2 + d_4V_y^4 + d_5V_y^3y + d_6V_y^2y^2 + d_7V_yy^3) V^2 \end{aligned}$$

Hydrodynamic profiles

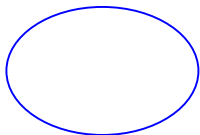
$$p(y) = p_0 \left[1 + \frac{6}{5} \left(\frac{mg}{T_0} \right)^2 y^2 \right] + \mathcal{O}(g^4)$$

$$u_z(y) = u_0 + \frac{\rho_0 g}{2\eta_0} y^2 + \mathcal{O}(g^3)$$

$$T(y) = T_0 \left[1 - \frac{\rho_0^2 g^2}{12\eta_0 \kappa_0 T_0} y^4 + \frac{1}{25} \frac{38 + 43\zeta_0^* + 17\zeta_0^{*2}}{(1 + \zeta_0^*)(2 + \zeta_0^*)} \left(\frac{mg}{T_0} \right)^2 y^2 \right] + \mathcal{O}(g^4)$$



NS terms



Extra terms

$$\zeta_0^* = \frac{\frac{5}{12}(1 - \alpha^2)}{\beta(\alpha) + \frac{5}{12}(1 - \alpha^2)}$$

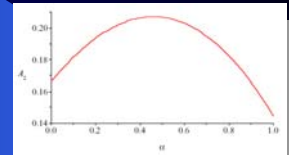
Non-monotonic temperature profile

$$T = T_0 \left[1 - A_4(\alpha) \left(\frac{g\lambda_0}{v_0^2} \right)^2 \left(\frac{y}{\lambda_0} \right)^4 + A_2(\alpha) \left(\frac{g\lambda_0}{v_0^2} \right)^2 \left(\frac{y}{\lambda_0} \right)^2 \right] + \mathcal{O}(g^4)$$

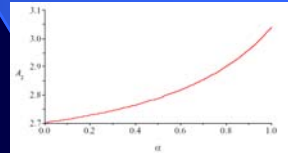
$$\eta, \kappa = \text{Boltzmann} \Rightarrow A_4(\alpha) = \frac{4}{1125\pi} (1 + \alpha)^2 (3 - \alpha) (49 - 33\alpha)$$

$$\beta(\alpha) = (1 + \alpha) \frac{2 + \alpha}{6} \Rightarrow A_2(\alpha) = \frac{4}{25} \frac{2719 - 2741\alpha + 706\alpha^2}{(7 - 4\alpha)(23 - 11\alpha)}$$

NS term

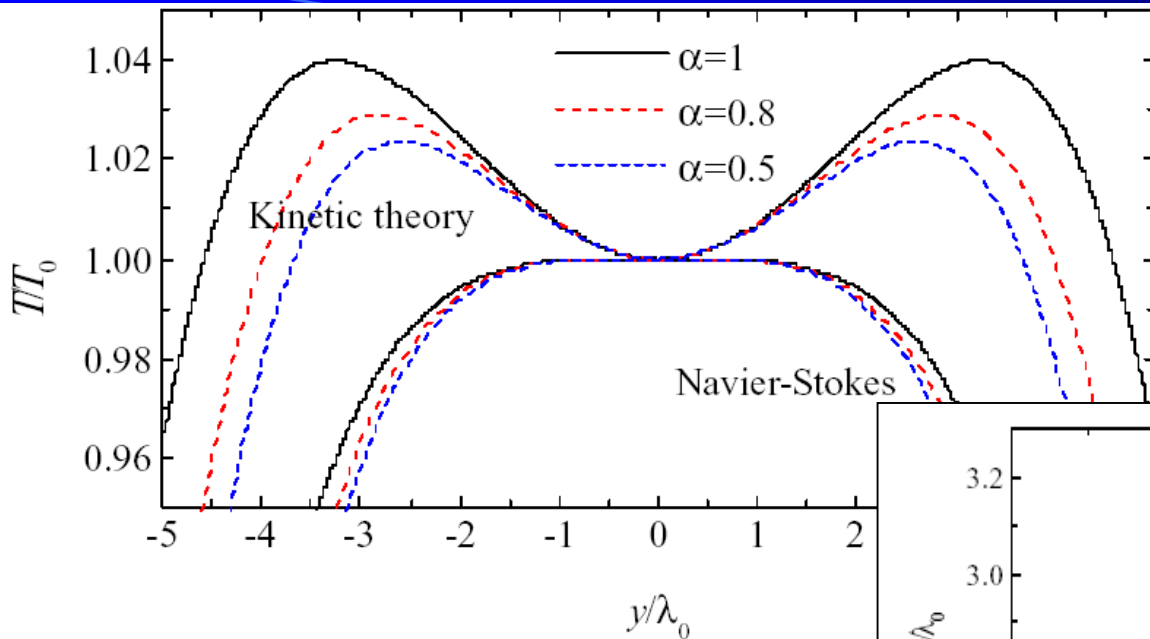


Extra term



(independent of g)
$$y_{\max} = \pm \lambda_0 \sqrt{\frac{A_2(\alpha)}{2A_4(\alpha)}}$$

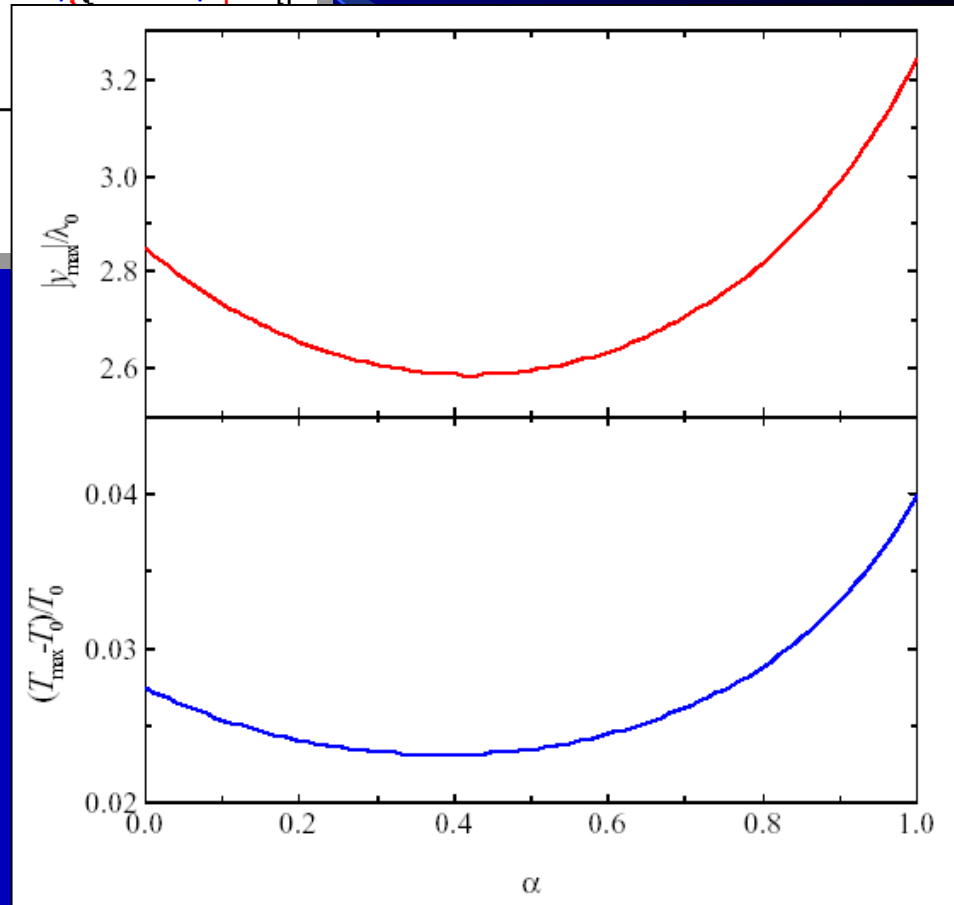
$$\frac{T_{\max} - T_0}{T_0} = \frac{A_2^2(\alpha)}{4A_4(\alpha)} \left(\frac{g\lambda_0}{v_0^2} \right)^2 + \mathcal{O}(g^4)$$



$$g\lambda_0/v_0^2=0.05$$

If $\alpha \gtrsim 0.4$, the bi-modal shape of $T(y)$ becomes (slightly) less pronounced as inelasticity increases.

However, the opposite behavior takes place if $\alpha \lesssim 0.4$.



Fluxes

$$P_{yz}(y) = -\rho_0 g y + \mathcal{O}(g^3)$$

$$P_{yy} = p_0 \left[1 - \frac{12}{25} \frac{102 + 87\zeta_0^* + 13\zeta_0^{*2}}{(1 + \zeta_0^*)(2 + \zeta_0^*)^2} \frac{\rho_0 \eta_0^2 g^2}{p_0^3} \right] + \mathcal{O}(g^4)$$

$$P_{zz}(y) = p_0 \left[1 + \frac{16}{25} \frac{82 + 67\zeta_0^* + 8\zeta_0^{*2}}{(1 + \zeta_0^*)(2 + \zeta_0^*)^2} \frac{\rho_0 \eta_0^2 g^2}{p_0^3} + \frac{14}{5} \left(\frac{mg}{T_0} \right)^2 y^2 \right] + \mathcal{O}(g^4)$$

Normal stress differences

$$q_y(y) = \frac{\rho_0^2 g^2}{3\eta_0} y^3 + \mathcal{O}(g^4)$$

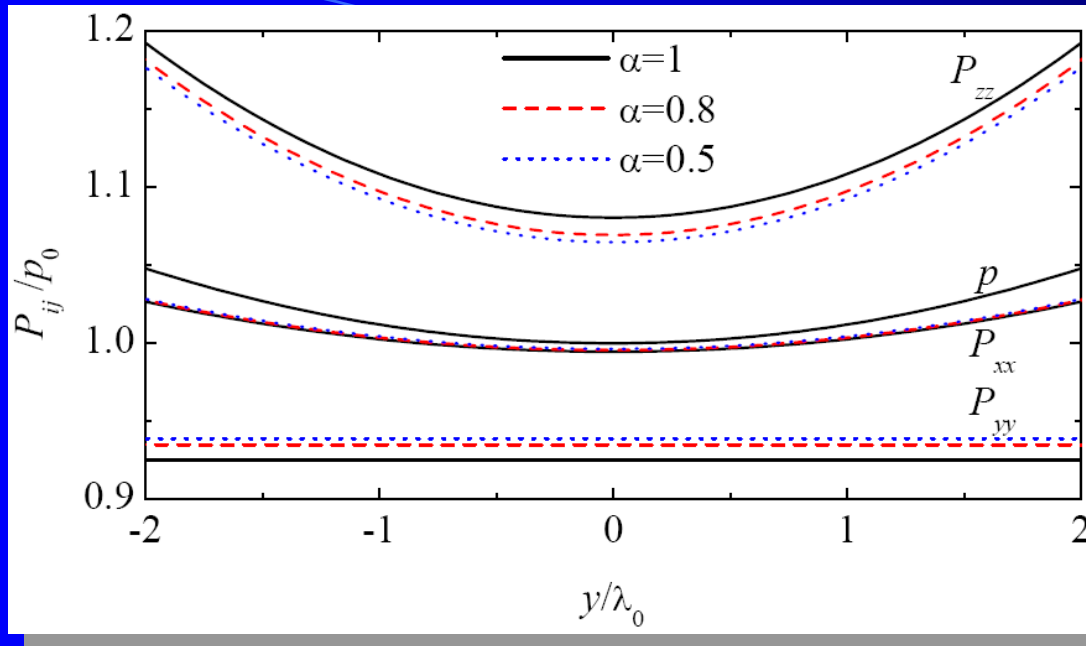
$$q_y = -\kappa \frac{\partial}{\partial y} \left(T + \frac{y_{\max}^2}{6} \nabla^2 T \right) + \mathcal{O}(g^4)$$

$$q_z = \frac{2}{5} m \kappa_0 g + \mathcal{O}(g^3)$$

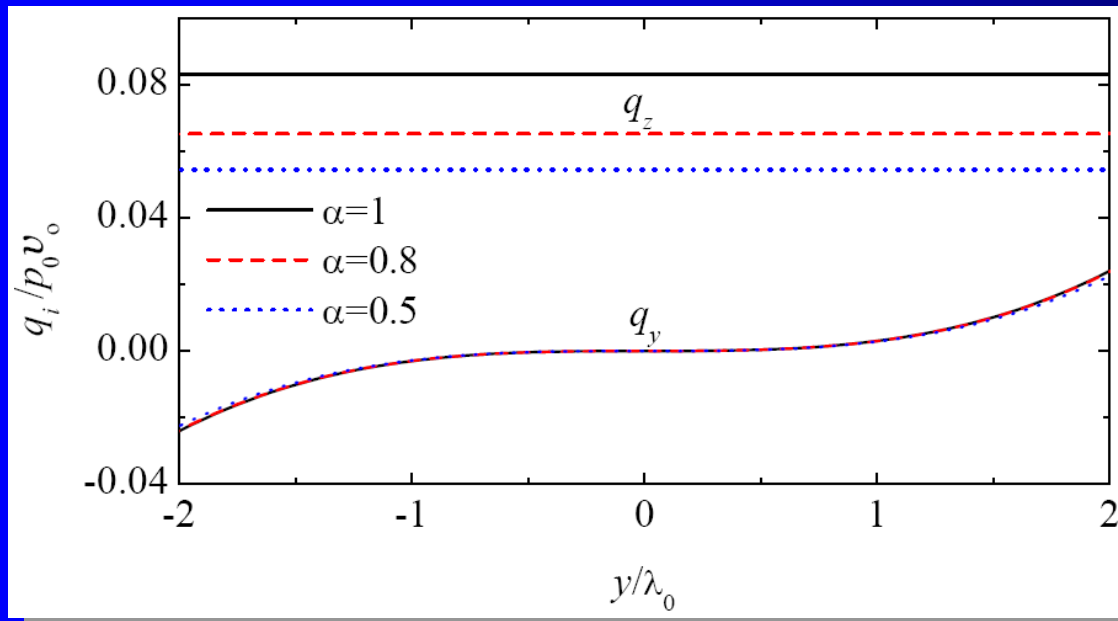
Longitudinal heat flux

Super-Burnett

$$g\lambda_0/v_0^2=0.05$$



$$P_{yy} < P_{xx} < p < P_{zz}$$



$$|q_y| < q_z$$

Conclusions (I)

- Gravity-driven Poiseuille flow exhibits interesting (and even counter-intuitive) non-Newtonian properties which are accessible to granular gases.
- Non-uniform hydrostatic pressure.
- Non-isotropic normal stresses.
- Heat flux component normal to the thermal gradient.

Conclusions (II)

- Bi-modal shape of the temperature profile:

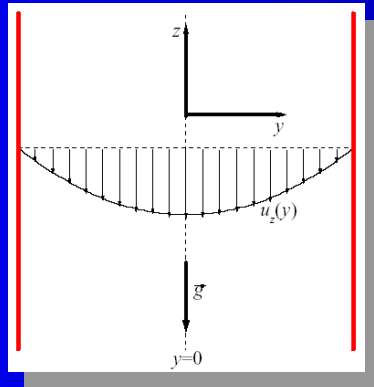
$$|y_{\max}| \approx 3 \text{ mfp}, (T_{\max} - T_0)/T_0 \approx 10 (g\lambda/v_{\text{th}}^2)^2.$$

- For moderate or small inelasticity ($\alpha \gtrsim 0.4$), the larger the inelasticity, the more pronounced the bi-modal temperature profile.

The reverse is true for large inelasticity ($\alpha \lesssim 0.4$).

- A similar influence of α on normal stress differences.
- Computer simulations (DSMC or MD) would be very welcome!

THANKS!



$$\left(-\frac{\zeta T}{2m} \frac{\partial^2}{\partial \mathbf{v}^2} - g \frac{\partial}{\partial v_z} + v_y \frac{\partial}{\partial y} \right) f = J[f, f]$$

